# **PHENIX Background Status**

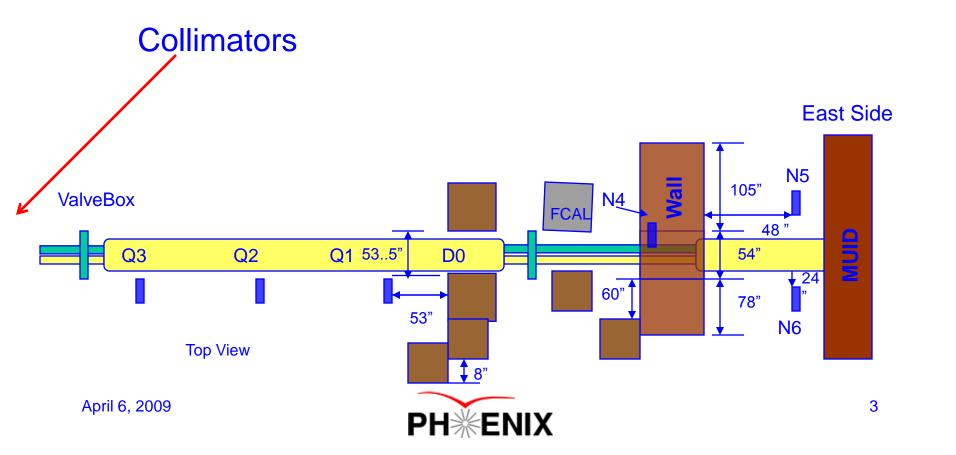
John Haggerty *Brookhaven National Laboratory* 

### The weekend background

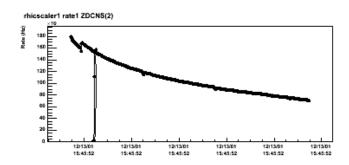
We decided to turn on the MUID and RPC starting Friday night when the background counters were higher than usual by a factor of ~2 (500 kHz in N3)

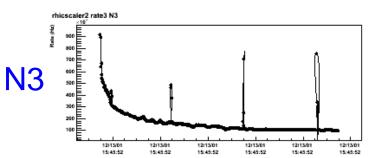
PHENIX history and lore is that we should wait until the counters outside Q1Q2Q3 are "low" which has historically been redefined (up) over runs

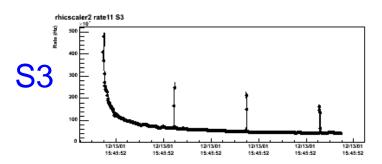
# Layout

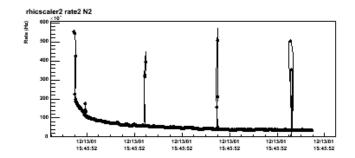


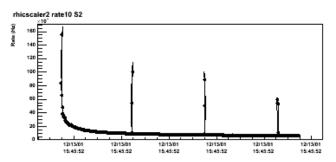
# **Background counters**





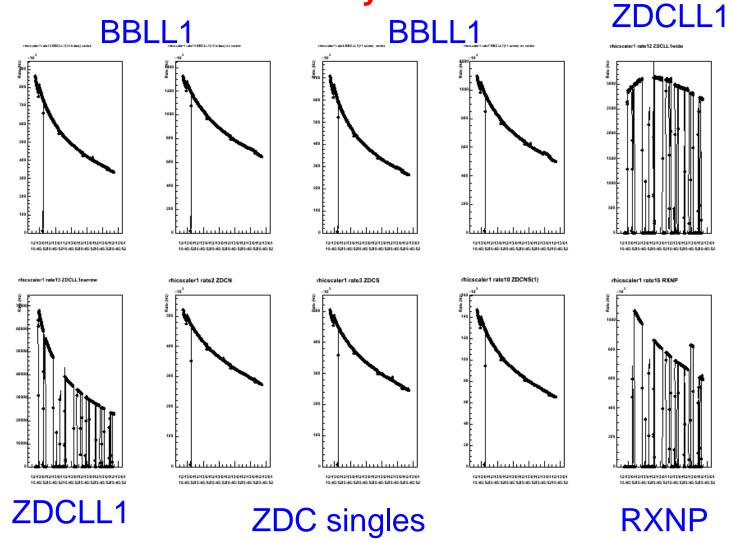




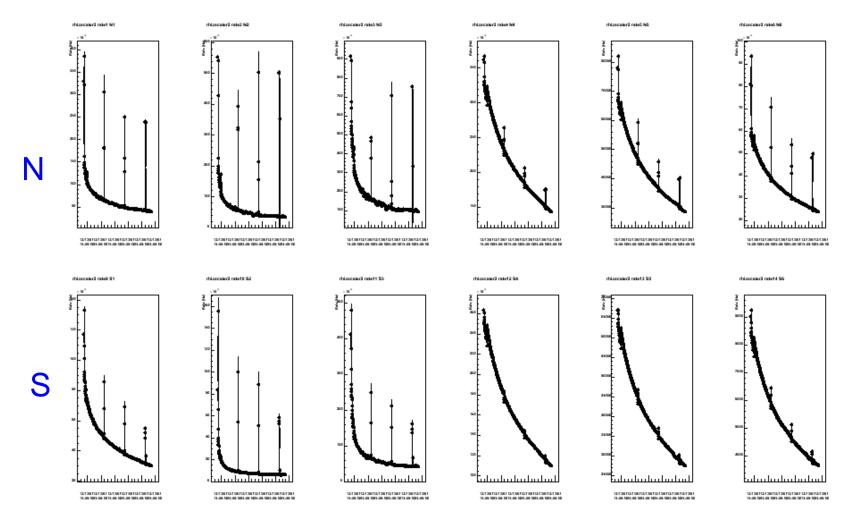




### **Luminosity scalers**

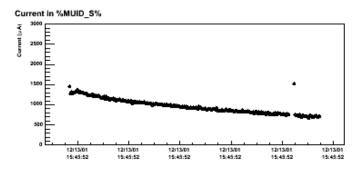


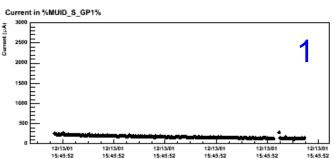
# All background counters

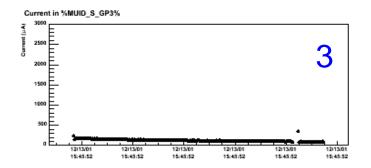


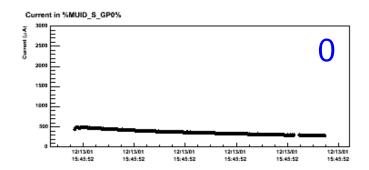


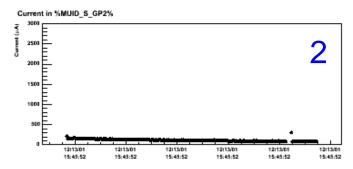
### **MUID.S Currents**

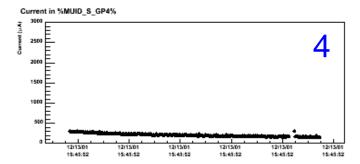






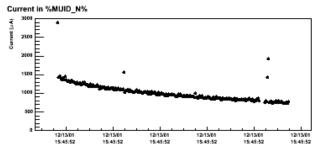


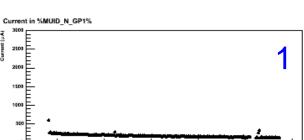






### **MUID.N Currents**





12/13/01

15:45:52

12/13/01

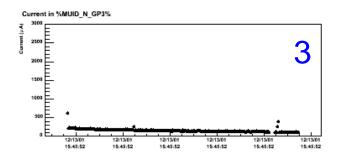
15:45:52

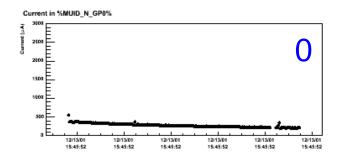
12:13/01

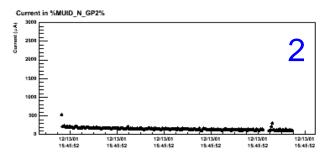
15:45:52

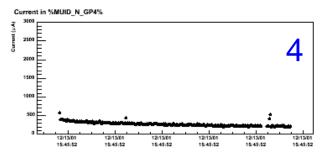
12:13/01

15:45:52











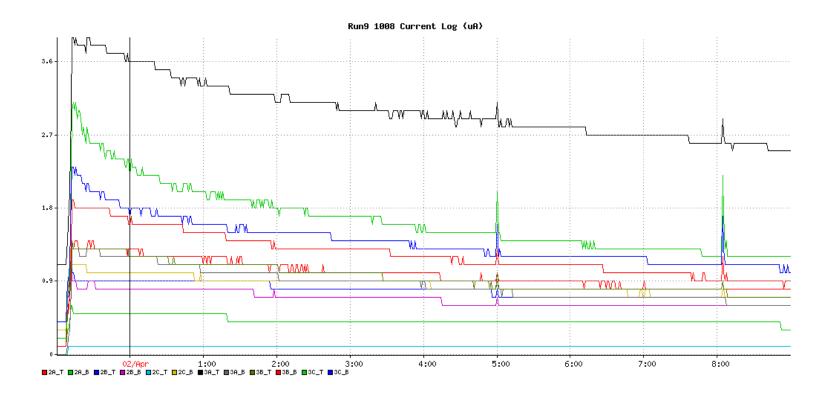
12/13/01

15:45:52

12/13/01

15:45:52

## RPC.S





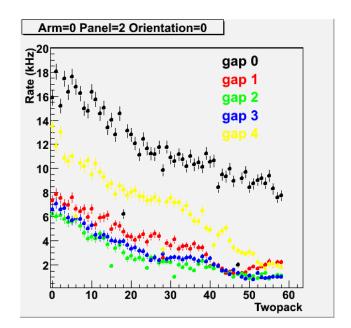
### **Trips**

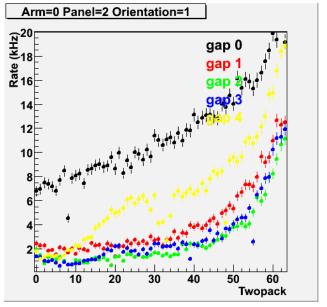
- DC 12 (see below)
- PC 20
- MUID 151 (only 5 before midnight)
- TEC 1

```
HV DC W S KS00-07 UV2 GBS |
                              278402 | 2009-04-02 05:38:19.716484-04 |
                                                                          0 | {1873466,-1262485488,-1263524392,9544888,83,6950439,-1208941832,4764032}
HV DC W S KS00-07 UV2 PBS |
                              278402 | 2009-04-02 05:38:19.772817-04 |
                                                                          0 | {0,0,-1208941848,4668387,2682868,-1262485488,-1263524392,-1208941848}
                                                                                                                                                        channel trip
                                                                          0 | {1873466,-1262485488,-1263524392,9544888,83,6950439,-1208941832,4764032}
                                                                                                                                                        channel trip
HV DC W S KS00-07 X2 GBS
                             278402 | 2009-04-02 05:38:19.802556-04 |
HV DC W S KS00-07 X2 PBS |
                             278402 | 2009-04-02 05:38:19.832521-04 |
                                                                         48 | {48,48,48,48,1280,1280,1280,1280}
                                                                                                                                                         channel trip
HV DC E S KS04-11 UV1 PBS |
                              278404 | 2009-04-02 06:51:32.79404-04 |
                                                                          1 | {512,16384,1,1,1,0,1,1}
                                                                                                                                                         channel trip
HV DC E S KS04-11 X1 BBS
                              278404 | 2009-04-02 06:51:33.494747-04 |
                                                                          1 | {156,0,-1229921512,4723652,165995600,1,1,4723630}
                                                                                                                                                        channel trip
                             278404 | 2009-04-02 06:51:33.62277-04 | 16384 | {1,1,1,1,6950439,1,4764032}
HV DC E S KS04-11 X1 GBS |
                                                                                                                                                         channel trip
                             278404 | 2009-04-02 06:51:33.853524-04 | 0 | {0,16384,1,1,1,0,1,1}
HV DC E S KS04-11 X1 PBS |
                                                                                                                                                         channel trip
HV DC E S KS12-19 UV1 GBS |
                              278404 | 2009-04-02 06:51:33.883717-04 | 16384 | {1,1,1,1,1,6950439,1,4764032}
                                                                                                                                                        channel trip
                              278404 | 2009-04-02 06:51:33.903497-04 | 0 | {0,16384,-1262483556,1,1,1,1,1}
HV DC E S KS12-19 UV1 PBS |
                                                                                                                                                         channel trip
HV DC E S KS12-19 X1 GBS
                              278404 | 2009-04-02 06:51:33.904028-04 | 16384 | {1,1,1,1,1,6950439,1,4764032}
                                                                                                                                                        | channel trip
HV DC E S KS12-19 X1 PBS
                              278404 | 2009-04-02 06:51:33.93351-04 | 0 | {-1,16384,-1262483556,1,1,1,1,1,1}
                                                                                                                                                       | channel trip
```

### **MUID** hit rates

- MUID hit rates can be used to estimate background rates
- Most reliable way is to take a completely unbiased trigger and count hits—clock triggers
- An example from Itaru
   Nakagawa shows
   increase in hits in gap 4
   (farthest from IP)







### Why we care

- Historically, the worry has been that varied current draw in the MUID would lead to varying muon trigger efficiency depending on beam conditions
- We are now trying to determine whether that's true or whether we might be able to live with it
- Of course, an increased hit rate in the MUTR and MUID leads to more confusion, probably reducing the resolution, but that's a fact of life at high luminosity

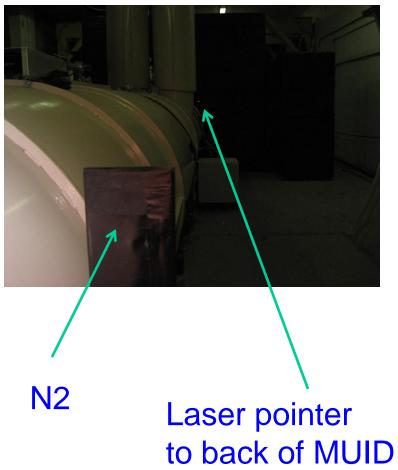
### **RPC**

 New detector, prototype in south tunnel (and IR)



### **Shielding holes**

- There are some holes in the shielding
- We should see whether we can plug them, particularly around the beam pipe



### Radiological survey

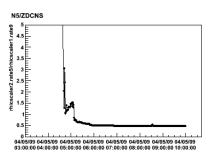
- I had HP do a radiological survey of the area around the collimators and quad triplet, and D0 magnet
- Collimators are pretty hot!
- Quadrupole activity is pretty uniform along the magnet, higher than it was

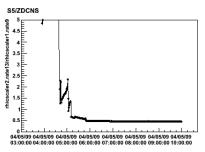
#### What to make of it?

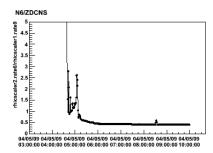
- Looks to me like 1,2,3 counters measure secondaries from the collimator which disappears fairly quickly as the beam is scraped
- Do the secondaries cause mischief in PHENIX IR?
   Not really any evidence for it

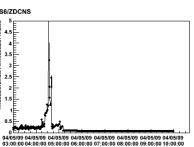
#### Scale to collision rate

- Started looking at background counters scaled to a collision rate (chose ZDCNS)
- Look only at counters 5 and 6 after the shielding



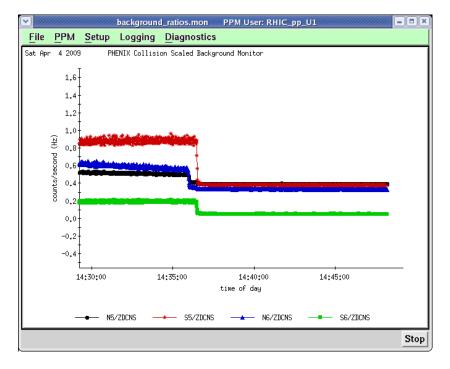






### **New background limits**

- Over the weekend, I started having the shift crews ramp up as soon as the 5,6 ratios to ZDCNS were about 0.5 as seen in background\_ratios.mon
- OK so far...



#### What next?

- If we can accumulate enough data with high background rates, we should be able to measure the efficiency as a function of rate
- We don't have good measurements of background in the region that will have the Si vertex detectors...
- I have come to think that much of what the 1,2,3 counters see is splash that we don't actually care about, but it may find shielding holes; should we just shield the quads and collimator?